



DOI: 10.1002/ange.201307851

## Visible-Light Photocatalytic Conversion of Carbon Monoxide to Methane by Nickel(II) Oxide\*\*

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Research in the area of "solar fuels" obtained photocatalytically by sunlight is actual and active because these fuels are sustainable and  $CO_2$ -neutral. Although  $H_2$  generation from  $H_2O$  is the most widely studied photocatalytic process for solar fuel production, the technical problems associated with  $H_2$  storage and fuel-cell efficiency make also attractive other chemicals obtained from  $CO_2$  such as  $CH_3OH$ ,  $CH_4$ , and  $CO.^{[1-7]}$ 

The production of solar fuels requires efficient photocatalysts with high visible-light photoresponse. While TiO<sub>2</sub> is the most widely studied photocatalyst, [8-10] this semiconductor shows activity only upon UV-light exposure and a considerable effort aims at developing TiO<sub>2</sub> photocatalysts with visible response.[11-15] In contrast, the use of p-type semiconductors absorbing visible light such as NiO, CuO<sub>r</sub>, FeO<sub>r</sub>, and CoO<sub>r</sub> has been less explored in photocatalysis, because of their presumably lower stability compared to TiO<sub>2</sub>. [16] In particular, NiO has been frequently employed as co-catalyst in combination with titanates or other photocatalysts for visible-light splitting of H<sub>2</sub>O rather than as active photocatalyst. In the present report, we describe that p-type semiconducting metal oxides, specifically NiO and FeOx, show high visible and sunlight photocatalytic activity for the reduction of CO by H<sub>2</sub>O vapor or H<sub>2</sub> leading to the formation of substantial amounts of CH<sub>4</sub>, with almost complete selectivity at full CO conversion. NiO can be reused without change in the XRD pattern, supporting its photostability. The novelty of our study derives from the consideration of two points: 1) the unprecedented room-temperature photocatalytic CO reduction by H<sub>2</sub>O or H<sub>2</sub> to CH<sub>4</sub>, and 2) the possibility of using sunlight and visible light to activate the process.

Preliminary experiments using  $TiO_2$  as photocatalyst showed that this n-type semiconductor leads mainly to oxidation of CO by  $H_2O$  vapor, forming quasi-stoichiometric amounts of  $CO_2$  and  $H_2$  [Eq. (1)],

$$CO + H_2O \xrightarrow[\text{photocatalyst}]{\text{hv}} CO_2 + H_2(\Delta G^\circ = +31.36 \text{ kJ mol}^{-1}) \eqno(1)$$

accompanied by detectable, amounts of  $CH_4$ , and traces of ethane. Formation of  $CO_2$  can be understood as arising from photooxidation of CO by positive holes generated in the charge separation on the semiconductor. The conversion of CO using  $TiO_2$  promoted by solar light was very small (see Table 1) in agreement with the well-documented low photocatalytic activity of  $TiO_2$  at solar light irradiation. [15]

**Table 1:** Photocatalytic activity for CO conversion, carbon content, and formation of  $H_2$ ,  $CO_2$  and  $CH_4$  in the presence of metal-oxide semi-conductors. [a]

Photocatalyst <sup>[b]</sup>	CO Conversion	C [%] <sup>[c]</sup>	Production [ $\mu$ mol g <sup>-1</sup> cat.]		
			$H_2$	$CO_2$	$CH_4^{[d]}$
TiO <sub>2</sub> <sup>[e]</sup>	3	0.1	260	650	4 (0.5)
NiO <sup>[e]</sup>	95	3.3	820	12200	1700
					(10.2)
CuO <sup>[e]</sup>	14	0.5	20	2900	270 (7.5)
$Co_3O_4^{[e]}$	91	1.2	1000	12300	600 (4.3)
$Co_2O_3^{[e]}$	93	3.1	990	12500	200 (1.3)
$Fe_3O_4^{[e]}$	59	0.1	3620	8200	900 (9.8)
$Fe_2O_3^{[e]}$	68	2.5	3760	10400	20 (0.2)
NiO <sup>[f]</sup>	96	2.2	_[g]	11 240	8387
					(39.1)
$Fe_3O_4^{[f]}$	49	3.6	_[g]	6590	733 (7.1)
NiO <sup>[h]</sup>	99.5	1.8	_[g]	711	17266 (87)
Fe <sub>3</sub> O <sub>4</sub> <sup>[h]</sup>	28	0.5	_[gj	943	580 (28.2)
NiO <sup>[i]</sup>	99.4	1.9	_[g]	450	20 799 (92)
NiO <sup>[j]</sup>	99.0	1.8	_[g]	87	16030 (97)
$NiO^{[k]}$	16	1.0	_[g]	1386	1379
					(38.3)

[a] Blank controls: 1) illumination in the absence of any solid, 2) contact to the solid and gas phase in the dark or 3) irradiation of the photocatalyst in N2 did not lead to any product, C mass balances were higher than 95%. [b] See the Supporting Information for the origin and characterization data of these commercially available photocatalysts. [c] Weight percentage of carbon on the solid after the photocatalytic experiment. [d] The number in brackets corresponds to the selectivity of the process to CH<sub>4</sub> defined as number of moles formed of CH<sub>4</sub> divided by moles of converted CO in percentage. [e] Irradiation conditions: Mixture of  $N_2$  (19.6 mmoles),  $H_2O$  (2.5 mmol), and CO (20%, 4.9 mmol), irradiation time 22 h, photocatalyst mass 250 mg, irradiation source solar simulator (1000 W m<sup>-2</sup>). [f] Reaction conditions: Mixture of CO (20%, 5 mmol), H<sub>2</sub> (5 mmol), Ar (1.1 mmol), and N<sub>2</sub> (13.9 or 0.125 mmol); photocatalyst mass 250 mg, reaction time 3 h, light source solar simulator. [g] Reagent. [h] Reaction conditions: Mixture of CO (20%, 5 mmol), H<sub>2</sub> (18.75 mmol), Ar (1.1 mmol), and N<sub>2</sub> (13.9 or 0.125 mmol); photocatalyst mass 250 mg, reaction time 3 h, light source solar simulator. [i] Second test under conditions indicated in [h] and activation of NiO at  $450\,^{\circ}\text{C}$  with air for 5 h. [j] Third test under conditions indicated in [h] and activation of NiO at 450°C with air for 5 h. [k] Irradiation with the solar simulator equipped with a cut-off filter ( $\lambda >$  400) under conditions indicated in [h].

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<sup>[\*\*]</sup> Financial support by the Spanish Ministry of Economy and Competitiveness (MINECO) through several actions (Severo Ochoa and grant number CTQ 2012-32315) is gratefully acknowledged.

Supporting information for this article is available on the WWW under http://dx.doi.org/10.1002/anie.201307851.



Then we tested other first-row transition-metal oxides that can absorb visible light, though they are generally considered much less efficient photocatalysts than TiO<sub>2</sub>. The results obtained for the p-type metal-oxide semiconductors are summarized in Table 1, whereas Figure S1 (see the Supporting Information) shows the temporal evolution of CH<sub>4</sub> upon irradiation in the presence of the p-type semiconductors using simulated solar light.

In contrast to experiments performed in the dark, where no products are formed (see footnote [a] in Table 1), or in contrast to the generation of CO2 using TiO2, p-type semiconducting metal oxides of Ni, Cu, Fe, and Co afford significant amounts of CH<sub>4</sub> in the gas phase upon sunlight irradiation. Equation (2) provides the presumed stoichiom-

$$CO + 2H_2O \rightarrow CH_4 + 1.5 O_2 \ (\Delta G^0 = 541 \text{ kJ mol}^{-1})$$
 (2)

etry of the CH<sub>4</sub> formation for the photocatalytic CO reduction by water using simulated solar light. While TiO<sub>2</sub> is a white powder, all the rest of the transition metals in Table 1 are colored, the visible absorption bands being responsible for the notable sunlight photoactivity (up to 33 times that of TiO<sub>2</sub>). Particularly notable is the performance of NiO and Fe<sub>3</sub>O<sub>4</sub> (in its most abundant  $\alpha$ -polymorph)<sup>[17]</sup> where the production of CH<sub>4</sub> is the highest (10.2% in the gas-phase mixture in the case of NiO). These differences in photocatalytic activity in the series of p-type metal-oxide semiconductors are unlikely to arise from differences in their surface area, since they show similar values (see Table S1). It is more reasonable to attribute the results to the intrinsic photocatalytic properties (redox potentials of electrons and holes, efficiency of charge separation and lifetime of charge separation among other parameters to be considered) of the semiconductor.

Figure 1 shows the temporal evolution of the CO conversion and the formation of CO<sub>2</sub>, CH<sub>4</sub>, and H<sub>2</sub> using NiO as photocatalyst and H<sub>2</sub>O as reducing agent under simulated sunlight irradiation. As can be seen in Figure 1, H<sub>2</sub> is formed

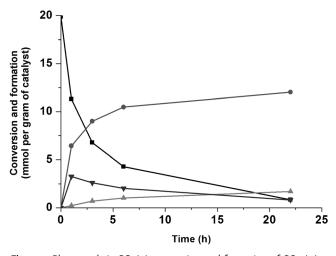


Figure 1. Photocatalytic CO ( $\blacksquare$ ) conversion and formation of CO<sub>2</sub> ( $\bullet$ ),  $CH_4$  ( $\blacktriangle$ ) and  $H_2$  ( $\blacktriangledown$ ). Reaction conditions: gas-phase irradiation of a moisture-saturated mixture of N<sub>2</sub> (80%) and CO (20%, 4.9 mmol), NiO 250 mg, irradiation source solar simulator (1000  $W \times m^{-2}$ ).

as primary, but unstable product and its proportion grows in the first hour of irradiation, but subsequently decreases gradually. CH<sub>4</sub> exhibits an induction period at initial times and is formed as a secondary product. These temporal profiles, particularly the fact that the rate of H<sub>2</sub> decrease is about one half that of the CH<sub>4</sub> growth, provide kinetic support to a CH<sub>4</sub> formation mechanism involving prior generation of H<sub>2</sub> from H<sub>2</sub>O vapor and subsequent reaction of H<sub>2</sub> to form CH<sub>4</sub>.

With regard to Table 1, it is worth commenting on the high percentage of carbon present on the solid photocatalyst, particularly in the cases of NiO, Co<sub>2</sub>O<sub>3</sub>, and Fe<sub>2</sub>O<sub>3</sub>. This carbonaceous residue adsorbed on the photocatalyst was characterized by Raman and X-ray photoelectron spectroscopy (XPS; see Figures S2 and S3), both spectra being compatible with the assignment of this residue to graphitic carbon generated in the photocatalytic process. The fresh photocatalysts do not show any carbon-related Raman peak. However, after irradiation, the characteristic 2D, G, and D bands of carbon appearing between 2500-2900 and at 1600 and 1350 cm<sup>-1</sup> were observed.

XPS of the photocatalyst after being used shows the presence of a peak at 285.12 eV that can be deconvoluted in three major components corresponding to sp<sup>2</sup> graphene C atoms (67%), sp<sup>2</sup> C atoms bonded to O (25%), and sp<sup>3</sup> C atoms (8%), respectively. The percentage of carbonaceous residue deposited on each material before and after irradiation was determined by combustion elemental analysis of the solid. The carbon content of the fresh photocatalysts was negligible, whereas it was necessary to consider the amount of carbon deposited on the photocatalyst after the reaction, together with the CH<sub>4</sub> and CO<sub>2</sub> formed in the gas phase to calculate satisfactorily the C mass balance (>95%) in all cases. This indicates that CO is converted not only into gasphase products, but it also forms carbon deposits on the solid. In agreement with the presence of carbon on the used photocatalyst, the colored metal oxides after the reaction become in some cases visually black (see the inset in Figure S2 for NiO).

The behavior observed for p-type semiconductors (formation of large CH<sub>4</sub> percentages in the gas phase and carbon residues on the solid) sharply contrasts with the photocatalytic activity of TiO<sub>2</sub> commented earlier [Eq. (1)]. Additional tests using CeO<sub>2</sub> as photocatalyst, another n-type semiconductor, also show the preferential formation of CO<sub>2</sub>. Sunlight conversion of CO using TiO<sub>2</sub> and CeO<sub>2</sub> is about 30 times lower than the conversion achieved by p-type metal oxides. This low activity of TiO<sub>2</sub> and CeO<sub>2</sub> at solar light irradiation is well-known and derives from their limited visible-light photoresponse of these wide-bandgap semiconductors. Besides the advantageous photoaction spectra of n-type semiconductors arising from the visible-light absorption of these materials, the different photocatalytic behavior of pand n-type semiconductors could derive from the higher conduction-band potential in p-type semiconductors (+0.8 V versus the normal hydrogen electrode, NHE, for Ni)[18] compared to  ${\rm TiO_2}$   $(-0.1\,{\rm V}$  versus NHE at pH 0)<sup>[19]</sup> and CeO<sub>2</sub> (-0.1 V).<sup>[12]</sup> A high conduction band potential favors the reduction of CO to C and CH4 with respect to the

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oxidation of CO to  $CO_2$  which is almost the sole process observed for  $TiO_2$  and  $CeO_2$ .

To establish if the photocatalytic reduction of CO to  $CH_4$  using p-type photocatalyst can be further favored using  $H_2$  instead of  $H_2O$  as reducing agent present in the system, additional experiments with simulated sunlight where NiO or  $Fe_3O_4$  were illuminated in contact with an atmosphere that contains CO and  $H_2$  were carried out.  $H_2$  could be obtained independently from steam reforming and the water-gas shift reaction or generated photocatalytically from water. The results obtained using two different  $CO/H_2$  molar ratios are summarized in Table 1 (footnotes [f] and [h]), whereas the corresponding time conversion plots are shown in Figure 2.

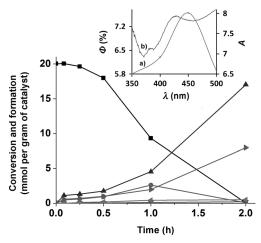


Figure 2. CO (■) conversion and formation of CO<sub>2</sub> (•) and CH<sub>4</sub> ( $\blacktriangle$ ), O<sub>2</sub> (•), ethane ( $\blacktriangleleft$ ) and propane ( $\blacktriangledown$ ). Reaction conditions: Mixture of CO (20%, 5 mmol), H<sub>2</sub> (18.75 mmol), Ar (1.1 mmol), and N<sub>2</sub> (13.9 or 0.125 mmol); photocatalyst mass 250 mg, reaction time 3 h, light source solar simulator, illumination source solar simulator (1000 W×m<sup>-2</sup>). The inset shows the apparent quantum yields ( $\Phi$ ) for the CH<sub>4</sub> formation (a) and the UV/Vis absorption spectrum of NiO (b; A = absorbance).

The photocatalytic  $CH_4$  formation increases considerably using  $H_2$  as reducing agent. The high CO conversions achieved using  $H_2$  allow the quantification of the  $O_2$  evolution, which shows a trend similar to the temporal profile of the  $CH_4$  evolution (Figure 2) according Equation (3).

$$CO + 2H_2 \rightarrow CH_4 + 0.5 O_2(\Delta G^0 = +35.6 \text{ kJ mol}^{-1})$$
 (3)

Both in the absence and presence of  $H_2$ , formation of  $CO_2$  is observed although with considerably lower yields when  $H_2$  is present (see Table 1 and Figure 2). This higher activity using  $H_2$  as reducing agent can be easily understood considering the high endoergonicity of the reaction using  $H_2O$  as reducing agent [Eq. (2)] and the more complex reaction mechanism using  $H_2O$  that requires prior  $H_2$  generation. The percentage of elemental carbon deposited on NiO is reduced by one half when the irradiation is carried out using  $H_2$  as reducing agent compared to  $H_2O$  vapor.

For the formation of  $CH_4$  two equivalents of  $H_2$  should be consumed. Thus, for those irradiation reactions where  $H_2$  and

CO are present in a 1:1 molar ratio, the generation of CH<sub>4</sub> reaches asymptotically a value that does not grow further when the reaction time is prolonged (Figure S4). When the irradiation is carried out in H<sub>2</sub> excess, the CH<sub>4</sub> production is higher, the consumption of CO is complete, and the percentage of CO<sub>2</sub> formed is considerably lower.

Comparison of the photocatalytic activity of NiO and  ${\rm Fe_3O_4}$  indicates that the former is more efficient for the  ${\rm CH_4}$  formation (Table 1). The highest  ${\rm CH_4}$  yield achieved herein was 17.26 mmol per gram of catalyst, corresponding to a reaction carried out using NiO as photocatalyst with a 3.5-fold excess of  ${\rm H_2}$ . Under these optimal conditions in excess of  ${\rm H_2}$  using simulated sunlight with a nominal power of  $1000~{\rm W}\times{\rm m}^{-2}$ , the yield of  ${\rm CO_2}$  is 24 times lower than the global amount of alkanes. This efficiency using simulated solar light is remarkable and without precedent.

To assess the visible-light photoactivity of NiO, the photoaction spectrum for the CH<sub>4</sub> generation was determined (see the inset in Figure 2). The overlay of the apparent quantum yield for the CH<sub>4</sub> generation based on the incident photon conversion (maximum  $\Phi_{\rm app}$  7.5% at 450 nm) with the NiO absorption band shows a notable coincidence, although the efficiency of the photocatalytic CH<sub>4</sub> generation at long wavelengths is much lower compared to the NiO absorption spectrum. This could be related to the higher efficiency of the ligand-to-metal electronic transition appearing around 410 nm with respect to the less efficient Ni<sup>2+</sup>-centered d–d absorption characteristic of transition-metal ions with incomplete d shell appearing at longer wavelengths in the visible region.

In cases in which H<sub>2</sub> is present together with CH<sub>4</sub>, significant amounts of ethane and propane are also formed. Formation of these light alkanes decreases somewhat the CH<sub>4</sub> selectivity. This behavior is reminiscent of the Fischer–Tropsch process, where a mixture of CO and H<sub>2</sub> gives rise to the formation of linear hydrocarbon molecules at elevated temperatures using Ni or Fe catalysts. It is, however, pertinent to remind that the current process is performed at temperatures below 36°C.

Formation of elemental carbon on the photocatalyst corresponds to CO reduction and this elemental carbon could act as an intermediate in the formation of CH4. In order to verify this possibility, a test was carried out in which CO diluted in N2 was submitted to irradiation until the NiO photocatalyst became visually black. At this time, the irradiation was stopped and the photoreactor outgassed to remove CO. Sampling of the photocatalyst revealed 1.7 wt % carbon content. Then, H<sub>2</sub> was admitted into the photoreactor and the irradiation continued. CH<sub>4</sub> evolution was observed in the gas phase in the absence of CO. This experiment can be easily interpreted considering that upon irradiation in the absence of H<sub>2</sub> carbon deposits are formed on the NiO surface. Subsequently, this carbon residue is able to react photocatalytically with H<sub>2</sub> to form CH<sub>4</sub>, regardless of the absence of CO. A control experiment following an identical protocol in the dark shows that no CH<sub>4</sub> is formed.

A mechanistic proposal for the photoreduction of CO by H<sub>2</sub> using a p-type photocatalyst is shown in Scheme 1. According to this proposal, CO can act as a hole-trapping



A) n-Type semiconductors

CO oxidation
$$\begin{array}{cccc}
H_2O & + & 2e^- & \longrightarrow & H_2 \\
CO & + & 2h^+ & \longrightarrow & CO_2
\end{array}$$

B) p-Type semiconductors

CO oxidation
$$\begin{bmatrix}
H_2O + 2e^{-} & & H_2 \\
CO + 2h^{+} & & CO_2
\end{bmatrix}$$

CO reduction 
$$CO + 4H^+ + 4e^- \longrightarrow CH_4 + 1/_2O_2$$
  $2H_2O + 4h^+ \longrightarrow O_2 + 4H^+$ 

**Scheme 1.** Contrasting photocatalytic behavior of n- and p-type semiconductors based on the product distribution observed.

agent, leading to the formation of  $CO_2$ . If another electron donor agent, such as  $H_2O$  or  $H_2$ , is also present, then these compounds will compete with CO for oxidation. CO can alternatively undergo reduction by the semiconductor conduction-band electrons, forming elemental C and  $O_2$ . This CO reduction process will be more favorable as the reduction potential of the conduction-band edge of the photocatalyst increases (case of p-type semiconductors), as has been observed in related precedents. [20] Thus, one of the main factors influencing the outcome of the photocatalytic CO reduction should be the position of the semiconductor conduction-band edge.

In agreement with this rationalization, a stoichiometric growth in the concentration of molecular  $O_2$  was detected when  $H_2$  is the reducing agent. When  $H_2O$  is used as electron donor, the lower formation of  $CH_4$  makes the quantification of  $O_2$  uncertain because of the air contamination. Elemental carbon is deposited on the surface of the photocatalyst as evidenced visually (inset in Figure S2) and determined by combustion chemical analysis (Table 1). Elemental carbon undergoes efficient photocatalytic  $H_2$  reduction, giving rise to  $CH_4$ .

To provide some support to the mechanistic proposal, additional experiments were carried out. A mixture of CO and  $H_2$  was irradiated in the presence of methanol as sacrificial electron donor agent, whereby total inhibition of the formation of  $CO_2$  was observed, while  $CH_4$  was still formed. This is in accordance with the proposed origin of  $CO_2$  arising from the photooxidation of CO on the surface of the semiconductor as indicated in Scheme 1.

The photocatalytic irradiation of CO and  $H_2$  using simulated sunlight was also performed in the presence of CH<sub>4</sub> and NiO as photocatalyst, and under these conditions ethane as primary irradiation product was observed, similar to the Fischer–Tropsch mechanisms and the reported photocatalytic CH<sub>4</sub> conversion.<sup>[21]</sup>

One of the main reasons frequently invoked to disregard p-type semiconducting metal oxides as photocatalysts is their presumed lack of stability and the occurrence of corrosion. Photocorrosion tests have been carried out mainly in liquid water. Since these conditions are different from those employed in the present study in the gas phase, we have verified the stability of the photocatalytic activity of NiO by performing three consecutive tests of the same sample. After each test, the NiO sample was submitted to calcination at 450 °C in air to effect the combustion of possible carbona-

ceous residue. As can be seen in Table 1, a minor decay in the CO conversion (from 99.5 to 99%) and even an increasing CH<sub>4</sub> selectivity (from 87 to 97%) were measured after the three tests. Moreover, XRD of the NiO sample before and after the three tests was also coincident, ruling out changes in the crystal structure of NiO. XPS, which is a specific surface analysis technique, did not show differences in the binding energy related to the Ni 2p peaks (Figure S5). This information indicates the stability of NiO as photocatalyst under the present conditions. A similar stability was observed by XPS for Fe<sub>3</sub>O<sub>4</sub> and Co<sub>3</sub>O<sub>4</sub> (Figures S6–S8)

The previous photocatalytic experiments were carried out using simulated sunlight containing about 5% UV radiation. The high photocatalytic activity in terms of CO conversion using NiO as catalyst compared to TiO<sub>2</sub> or CeO<sub>2</sub> has been attributed to the visible-light photoresponse of NiO in accordance to the photocation spectrum. Additional experiments on the photocatalytic reduction of CO to CH<sub>4</sub> by H<sub>2</sub> using NiO were also carried out under visible-light irradiation using a filter that cuts off radiation wavelengths shorter than 400 nm (see the spectrum of the cut-off filter in Figure S9). The results of the visible-light photocatalytic activity of NiO (see Table 1) show that although a large percentage of the catalytic activity of NiO is related to the UV region, NiO is still active under visible-light illumination.

In conclusion, here we have shown that p-type transitionmetal oxide semiconductors are efficient photocatalysts for the visible-light or solar-light, room-temperature reduction of CO to CH<sub>4</sub>, provided that H<sub>2</sub>O vapor or H<sub>2</sub> are present as reducing agents. This photochemical behavior contrasts with that of TiO<sub>2</sub> and CeO<sub>2</sub> (n-type semiconductors) that promote under the same conditions highly selectively the oxidation of CO to CO<sub>2</sub>. Formation of elemental carbon deposited on the photocatalyst surface is also observed. NiO was the most efficient photocatalyst among the series of first-raw transition-metal oxide semiconductors. Under optimal conditions using simulated sunlight and molecular H<sub>2</sub> as reducing agent, a selectivity towards the formation of CH<sub>4</sub> of 97% at complete CO conversion was obtained. Our results open new opportunities for the photocatalytic production of solar fuels at high conversion and selectivity.

Received: September 6, 2013 Published online: October 24, 2013

**Keywords:** methane formation  $\cdot$  metal oxides  $\cdot$  photocatalysis  $\cdot$  semiconductors  $\cdot$  solar fuels

<sup>[1]</sup> G. R. Dey, J. Nat. Gas Chem. 2007, 16, 217.

<sup>[2]</sup> S. Navalón, A. Dhakshinamoorthy, M. Álvaro, H. Garcia, ChemSusChem 2013, 6, 562.

<sup>[3]</sup> J.-S. Hwanga, J.-S. Changa, S.-E. Parkb, K. Ikeuec, M. Anpo, Top. Catal. 2005, 35, 311.

<sup>[4]</sup> T. Inoue, A. Fujishima, S. Konishi, K. Honda, *Nature* **1979**, 277, 637.

<sup>[5]</sup> K. Koci, L. Obalová, L. Matejová, D. Plachá, Z. Lacný, J. Jirkovský, O. Solcová, Appl. Catal. B 2009, 89, 494.

<sup>[6]</sup> W. Leitner, Angew. Chem. 1995, 107, 2391; Angew. Chem. Int. Ed. Engl. 1995, 34, 2207.



- [7] K. Mori, H. Yamashita, M. Anpo, RSC Adv. 2012, DOI: 10.1039/ c2ra01332k.
- [8] M. A. Fox, M. T. Dulay, Chem. Rev. 1993, 93, 341.
- [9] M. R. Hoffmann, S. T. Martin, W. Y. Choi, D. W. Bahnemann, Chem. Rev. 1995, 95, 69.
- [10] D. W. Bahnemann, Res. Chem. Intermed. 2000, 26, 207.
- [11] A. Primo, A. Corma, H. Garcia, Phys. Chem. Chem. Phys. 2011, 13, 886.
- [12] A. Primo, T. Marino, A. Corma, R. Molinari, H. Garcia, J. Am. Chem. Soc. 2011, 133, 6930.
- [13] S. Sakthivel, M. Janczarek, H. Kisch, J. Phys. Chem. B 2004, 108,
- [14] S. Sakthivel, M. V. Shankar, M. Palanichamy, B. Arabindoo, D. W. Bahnemann, V. Murugesan, Water Res. 2004, 38, 3001.

- [15] C. G. Silva, R. Juarez, T. Marino, R. Molinari, H. Garcia, J. Am. Chem. Soc. 2011, 133, 595.
- [16] J. M. Herrmann, ACS Symp. Ser. 1986, 298, 200.
- [17] G. Carraro, C. Maccato, A. Gasparotto, T. Montini, S. Turner, O. I. Lebedev, V. Gombac, G. Adami, G. Van Tendeloo, D. Barreca, P. Fornasiero, Adv. Funct. Mater. 2013, n/a.
- [18] M. P. Dare-Edwards, J. B. Goodenough, A. Hamnett, N. D. Nicholson, J. Chem. Soc. Faraday Trans. 2 1981, 77, 643.
- [19] M. Grätzel, Nature 2001, 414, 338.
- [20] Y. P. Xie, G. Liu, L. Yin, H.-M. Cheng, J. Mater. Chem. 2012, 22,
- [21] L. Yuliati, H. Yoshida, Chem. Soc. Rev. 2008, 37, 1592.

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